

Basics of Magnetohydrodynamics (MHD)

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What is MHD?

- MHD = fluid description of plasmas

Plasma physics

- Ionized gases
 - Conductivity $\sigma > 0$
- Macroscopically neutral (if $\mathbf{E}_{\text{ext}} = 0$):
 - Debye shielding generates Yukawa potentials:

$$\Phi(r) = -\frac{q}{r} \exp\left(-\sqrt{Z+1} \frac{r}{\lambda_D}\right); \text{ where } \lambda_D := \sqrt{\frac{k_B T}{4\pi \bar{n}_e e^2}}.$$

- Coulomb interaction effectively short-ranged
- Magnetic fields \mathbf{B} from small drift velocities $v_{\text{drift}} = |v_e - v_i|$
 - Can be quite large for high σ

Hydrodynamics

- Fluid description:
 - Many random, short-ranged interactions
 - Scale hierarchy: $\lambda \ll L \ll L_{\text{sys}}$
 - Collisional invariants: N, p, E

Kinetic theory:
Liouville/Boltzmann eq.:
$$\frac{df}{dt}(t, \mathbf{x}, \mathbf{p}) = C[f]$$

- Macroscopic conservation laws fixing the fluid variables ρ, P, \mathbf{v}

MHD = Hydrodynamics applied to plasma

- Main difference to neutral gas: Coulomb interaction
 - Electrons & ions form two separate, coupled fluids
 - *But*: Single-fluid behavior for negligible charge separation:

$$\lambda_D \ll L$$

- **B**-field adds contrib. \mathcal{M}_{ij} to the fluid stress-energy tensor T_{ij}
(or: Lorentz force into Vlasov equation)

MHD = Hydrodynamics applied to plasma

- Dynamical variables and evolution laws in MHD:
 - ρ : Continuity equation
 - \mathbf{v} : Euler (or Navier-Stokes) equation (+ Lorentz force)
 - P : Energy conservation equation (+ Ohmic heat gain term)
 - \mathbf{B} : Maxwell's equations

The induction equation

- Assumptions:
 - Non-relativistic plasma (linearize in $|\mathbf{v}|/c$)
 - High, spatially constant conductivity ($\sigma \gg T^{-1}$)
 - Ohm's law
- Evolution of \mathbf{B} -field then follows from Maxwell's equations:

$$\begin{aligned}\nabla \cdot \mathbf{E} &= 4\pi q & \nabla \times \mathbf{E} &= -\partial_{ct}\mathbf{B} \\ \nabla \cdot \mathbf{B} &= 0 & \nabla \times \mathbf{B} &= \frac{4\pi}{c}\mathbf{j} + \underbrace{\partial_{ct}\mathbf{E}}_{\sim \mathcal{O}(|v|^2/c^2) \cdot \mathcal{O}(\frac{B}{L})},\end{aligned}$$

- and implies the induction equation:

The induction equation

- Induction equation:

$$\frac{\partial \mathbf{B}}{\partial t} = \underbrace{\frac{c^2}{4\pi\sigma} \Delta \mathbf{B}}_{\text{diffusion}} + \underbrace{\nabla \times (\mathbf{v} \times \mathbf{B})}_{\text{advection}}$$

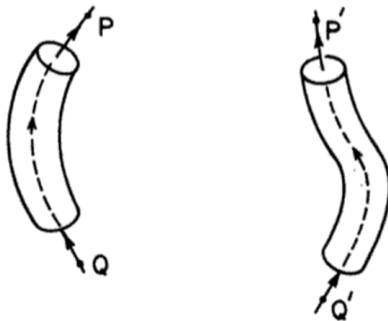
- Magnetic Reynolds number:

$$\mathcal{R}_M := \frac{\text{advection}}{\text{diffusion}} \sim \frac{v|\mathbf{B}|}{L} \bigg/ \frac{c^2}{4\pi\sigma} \frac{|\mathbf{B}|}{L^2} = \frac{4\pi\sigma Lv}{c^2}$$

- Diffusivity limit: $\mathcal{R}_M \ll 1$ (typical for lab situations)
- Ideal MHD limit: $\mathcal{R}_M \gg 1$ (typical for astrophys. systems)

Ideal MHD limit

- Perfect conductivity: $\sigma \rightarrow \infty$
 - $\frac{\partial \mathbf{B}}{\partial t} \approx \nabla \times (\mathbf{v} \times \mathbf{B})$
 - Field lines "frozen" into plasma (Alfvén's theorem of flux-freezing)
 - *i.e.* fluid elements connected by field line remain connected:



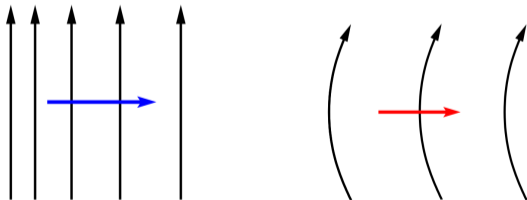
(Source: Choudhuri: The Physics of Fluids and Plasmas)

Magnetic field backreaction on plasma

- Euler equation modified by Lorentz force:

$$\begin{aligned} \rho \frac{d\mathbf{v}}{dt} &= -\nabla P + \frac{1}{c}(\mathbf{j} \times \mathbf{B}) \\ &\stackrel{\text{Ampère's law}}{=} -\nabla \left(P + \underbrace{\frac{\mathbf{B}^2}{8\pi}}_{\text{Mag. pressure}} \right) + \underbrace{\frac{1}{4\pi}(\mathbf{B} \cdot \nabla)\mathbf{B}}_{\text{Mag. tension}} \end{aligned}$$

- \Rightarrow Field lines resist to being squeezed or bent:



(Source: Nick Murphy, Lecture at Harvard-Smithsonian Center for Astrophysics)

Summary

- MHD = Dynamic theory of magnetized fluids
- Single-fluid approximation for tight Coulomb coupling
 - Induction equation fixes \mathbf{B}
 - Ideal MHD limit $\sigma \rightarrow \infty$: Advection only (Flux-freezing)
 - Backreaction of \mathbf{B} on fluid quantities ρ, \mathbf{v}, P via Lorentz-force